

A new index for the volcanic explosion scales using impulsive ionospheric disturbances

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Abstract: Using the total ionospheric electron content (TEC) data from ground-based global satellite navigation system (GNSS) receivers in Japan, we compared ionospheric responses to five explosive volcanic eruptions 2004-2015 of the Asama, Shin-Moe, Sakura-jima, and Kuchinoerabu-jima volcanoes. The TEC records show N-shaped disturbances with a period ~80 seconds propagating outward with the acoustic wave speed of ~0.8 km/s from the volcanoes. The amplitudes of these TEC disturbances are a few percent of the background absolute TEC, and we propose to use such relative amplitudes as a new index for the intensity of volcanic explosions.

Keywords: Ionospheric disturbance; Asama; Shin-Moe; Sakurajima; Kuchinoerabujima; volcanic explosion, index

1. Introduction

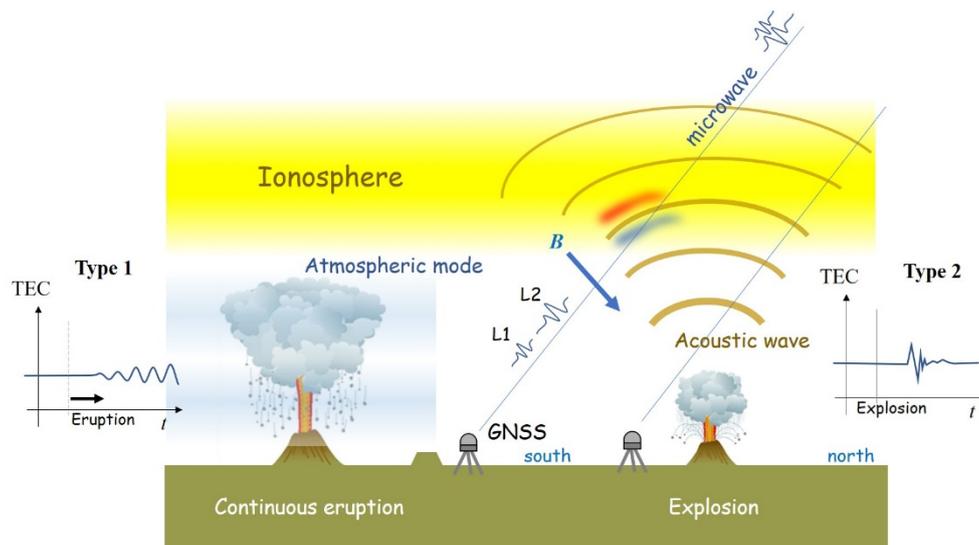
The Earth's ionosphere ranges from 60 to 1,000 km in altitude and is characterized by large number of free electrons. Ionospheric conditions are controlled by solar radiation and often disturbed by geomagnetic activities. In addition to such disturbances caused by space weather, the ionosphere is disturbed by events below (Blanc, 1985), such as earthquakes (Heki, 2020), tsunami (Occhipinti et al., 2013), and volcanic eruptions.

Ionospheric total electron content (TEC) can be easily measured by comparing the phases of two microwave carriers from global navigation satellite system (GNSS) satellites, such as Global Positioning System (GPS) (e.g. Hofmann-Wellenhof et al., 2008). Ground GNSS networks have been deployed to monitor crustal movements, and these networks were found useful to study ionospheric disturbances by volcanic eruptions. There are two types of ionospheric TEC responses to various types of volcanic eruptions.

The first type is the long-lasting harmonic TEC oscillations (Figure 1). It was found after the 13 July 2003 eruption of the Soufrière Hills volcano, Montserrat, in the Lesser Antilles (Dautermann et al., 2003a, b), and after the February 2014 eruption of the Kelud volcano, eastern Java Island, Indonesia (Nakashima et al., 2016). They reported that harmonic oscillations caused by atmospheric resonance excited by the Plinian eruption of the Kelud volcano lasted for ~ 2.5 hours after the eruption started. Shults et al. (2016) found similar TEC oscillations after the 2015 April Plinian eruption of the Calbuco volcano, Chile. Cahyadi et al. (2020) also found such harmonic TEC oscillations lasting ~20 minutes following the

43 2010 November 5 eruption of the Merapi volcano, central Java Island. They suggested that
44 the TEC oscillation amplitudes relative to background TEC represent the mass eruption rate,
45 and the products of such amplitudes and the duration provides an index for the total amount
46 of the ejecta.

47 The second type of disturbances occur 8-10 minutes after volcanic explosions by acoustic
48 waves propagating upward from the surface to ionospheric F region (Figure 1). They make
49 short-term N-shaped impulsive TEC responses as Heki (2006) observed with the GPS-TEC
50 method after the Vulcanian explosive eruption of the Asama volcano, Central Japan, on
51 September 1, 2004. In this work, we focus on the impulsive ionospheric responses to this
52 type of volcanic explosions.



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54 **Figure 1.** Ionospheric disturbance caused by continuous (Type 1 left) and explosive (Type 2 right)
55 volcanic eruptions can be detected by differential ionospheric delays of microwave signals of two
56 carrier frequencies (L1 and L2) from GNSS satellites. Strong continuous eruptions sometimes excite
57 atmospheric modes and long-term oscillatory disturbances in ionosphere. For explosive eruptions, we
58 often find short-term impulsive disturbances in ionosphere 8-10 minutes after eruptions, the time
59 required for acoustic waves to reach the ionospheric F region. The acoustic wave makes electron
60 density anomalies (pairs of positive and negative anomalies as shown with red and blue colors in the
61 figure) on the southern side of the volcano (for northern hemisphere cases).

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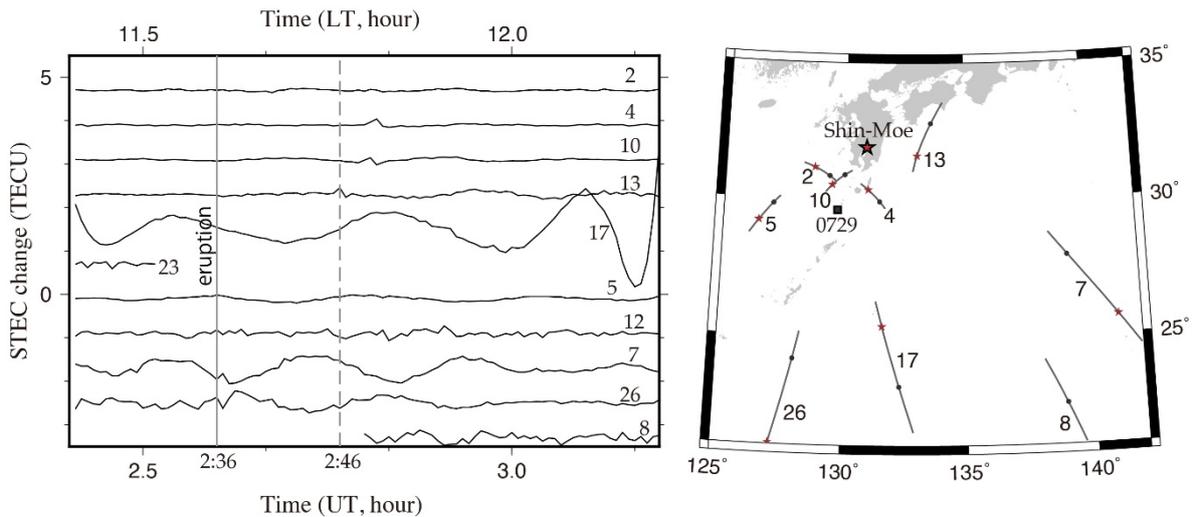
63 Intensity of a volcanic explosion has been studied by atmospheric pressure changes
64 associated with the airwave (infrasound) generated by the eruption (e.g. Matoza et al., 2019).
65 However, geometric settings of such sensors relative to volcanoes are diverse, and amplitudes
66 of such airwaves are difficult to serve as a universal index to describe the explosion intensity.
67 In this study, we explore the possibility to use the amplitude of ionospheric disturbance that
68 occur ~10 minutes after a large explosion as the new index. For this purpose, we compare
69 ionospheric TEC responses to five recent explosive volcanic eruptions of four volcanoes in
70 Japan 2004-2015 comparing the GNSS-TEC data from GEONET (GNSS Earth Observation
71 Network), a dense GNSS network in Japan.

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73 **2. GNSS data**

74 We calculated TEC by multiplying a certain factor to the phase difference of the
 75 microwave signals in two frequencies, L1 (~1.5 GHz) and L2 (~1.2 GHz), from GNSS
 76 receivers in the Japanese GEONET. TEC indicates number of electrons integrated along the
 77 line-of-sight (LoS) connecting the ground stations and GNSS satellites. We often represent
 78 the point of observation using the latitude and longitude of ionospheric piercing point (IPP),
 79 the point of intersection of LoS with the hypothetical thin layer assumed at the altitude of the
 80 highest electron density (~300 km). We plot their surface projections, often called sub-
 81 ionospheric points (SIP), on the map. In this study, we used only GPS satellites. The details
 82 of the GNSS-TEC technique to study lithospheric phenomena are available in the review
 83 article by Heki (2020).

84 We downloaded the raw GNSS data on the days of the eruptions from terras.gsi.go.jp. To
 85 remove the integer ambiguity of phase observables, the TEC obtained from carrier phases
 86 are adjusted to those from pseudo-ranges without such ambiguities. Such slant TEC (STEC)
 87 values are often converted to vertical TEC (VTEC) after removing inter-frequency biases in
 88 GNSS receivers and satellites. Here, however, we use STEC throughout the study in order to
 89 capture TEC signatures from LoS penetrating the wavefront with shallow angles (Figure 1).
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 92 **Figure 2.** STEC changes observed at the GNSS station 0729 (square in the map) over 2.4-3.2 UT,
 93 February 11, 2011. An explosive eruption of the Shin-Moe volcano (star in the map) occurred at 2:36
 94 UT (solid vertical line), and small ionospheric observations are seen to occur in signals with GPS
 95 satellites 4, 10, and 13 around 10 minutes after the eruption (dashed vertical line). In the map, we show
 96 the trajectory of SIPs with solid circles and red stars indicating the 3:00 UT and 2:36 UT, respectively.
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98 Interaction of the electron movement with geomagnetic fields allows us to observe such
 99 TEC disturbances from stations located to the south of volcanoes (e.g. see Rolland et al.,
 100 2013). Figure 2 shows an example of the STEC change time series before and after the 2011
 101 February 11 eruption of the Shin-Moe volcano in Kyushu, SW Japan, observed at the station
 102 0729 located in a small island to the south of Kyushu. We isolated the short-term signals by
 103 fitting the polynomials (with degree 7) to STEC time series and showing residuals from such
 104 reference curves. Small pulses occurring ~10 minutes after the eruption (Satellites 4, 10, 13)

105 are caused by acoustic waves propagating from the volcano to the ionosphere. Incessant small
106 fluctuations of TEC observed by Satellites 7, 8, 12, 17, and 26 are natural variabilities of
107 TEC intrinsic for low elevation satellites.

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109 **3. Five volcanic explosions of four volcanoes in Japan**

110 We selected five recent explosive volcanic eruptions in Japan with clear TEC disturbance
111 signals. The Asama volcano, central Japan, started eruptive activity at 11:02 UT on
112 September 1, 2004, with a Vulcanian explosion associated with strong airwaves (Nakada et
113 al., 2005). For this eruption, there have been reports of ionospheric disturbance using GNSS-
114 TEC by Heki (2006) and by HF-Doppler measurements by Chonan et al. (2018). Plume
115 height was unknown due to cloudy weather, and they detected the atmospheric pressure
116 change exceeding 205 Pa at a sensor located 8 km to the south (Yokoo et al., 2005).

117 Sakura-jima is an active stratovolcano in the Kagoshima Prefecture, Kyushu, and is one of
118 the most active volcanoes in SW Japan. Since 2009, several hundreds of explosions occur
119 every year in the volcano. There were 125 eruptions in 2009 October, in which 101 were
120 explosive. The explosion of Minamidake, Sakurajima, at 07:45 UT on October 3, 2009, was
121 one of the strongest eruptions in its activity since 2009. Plume reached the height ~3,000 m
122 above the caldera rim, and the atmospheric pressure change exceeding 294.5 Pa was detected
123 at a sensor 5 km southeastward. At another observatory, 11 km to the west of the vent,
124 pressure change of 74 Pa was observed (JMA, 2010).

125 Shin-Moe Volcano is also located in the Kagoshima Prefecture, Kyushu. We studied two
126 explosive eruptions in 2011. In the first eruption (Jan. 31 22:54 UT), the plume reached the
127 height of 2,000 m above the caldera rim, and the atmospheric pressure change exceeding
128 458.5 Pa was observed at a sensor 2.6 km southwestward. In the second eruption (Feb. 11,
129 02:36 UT), the plume reached the height of 2,500 m above the caldera rim, and the
130 atmospheric pressure change exceeding 244.3 Pa was observed at the same sensor (JMA,
131 2013).

132 The last volcanic eruption was Kuchinoerabu-jima volcano, located at a tiny island
133 Kuchinoerabu-jima ~100 km to the south of Kyushu. The eruption occurred on 3 January
134 2015 (00:59 UT). The plume height was 9,000 m, and pyroclastic flow reached the ocean.
135 An atmospheric pressure of 62.2 Pa was observed at a sensor located 2.3 km northeastward
136 (JMA, 2015). Nakashima (2018) studied ionospheric disturbances caused by this eruption by
137 using 1 Hz high-rate GNSS data.

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139 **4. Comparison of ionospheric disturbances by the five volcanic explosions**

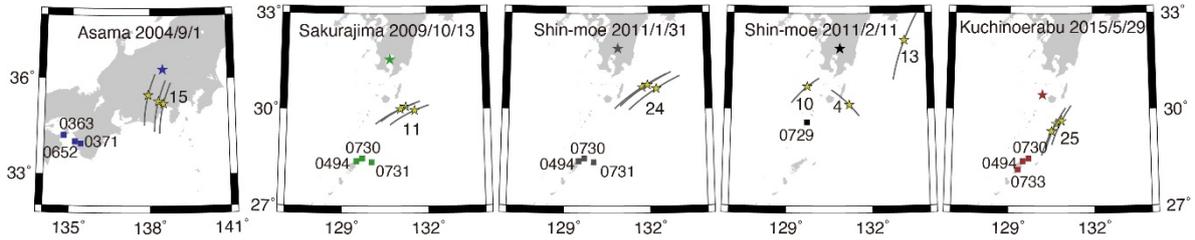
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141 Figure 3 compares the impulsive TEC changes with periods of 1-2 minutes, ~10 minutes
142 after the explosion. They are all STEC and the time series show the residual from the best-fit
143 polynomials with degrees 7-9. We also see faint harmonic oscillations (similar to Type 1
144 disturbance in Figure 1) sometimes follow the eruptions, e.g. the 2015 Kuchinoerabu-jima
145 eruption. Here, we focus on the N-shaped TEC disturbances.

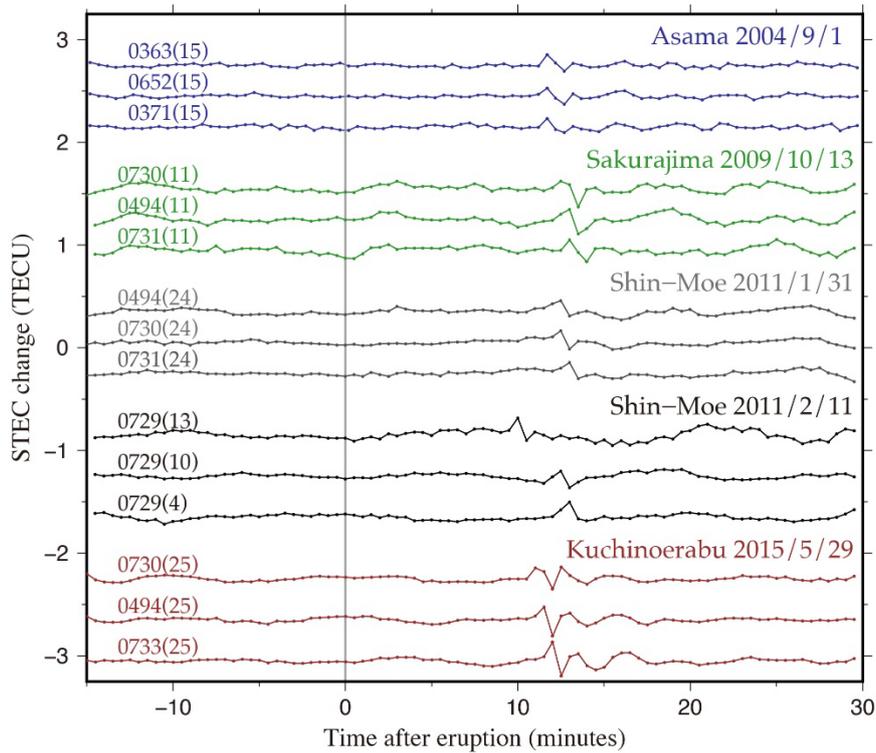
146 Strong disturbances can be seen only from GNSS stations located to the south of the
147 stations due to the interaction with geomagnetic fields. Therefore, we could not fully take

148 advantage of the dense GNSS network because the Sakura-jima, Shin-Moe and
 149 Kuchinoerabu-jima volcanoes are all located in southern Kyushu and GNSS stations are
 150 sparse to their south.

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155 **Figure 3.** (top) Geometry of volcanoes (large stars), SIP tracks (gray curves) and SIP positions at the
 156 time of the eruptions (small yellow stars), and GNSS stations (squares). We show the STEC time
 157 series for the three pairs of stations and satellites (three satellites from one station for the second Shin-
 158 Moe eruption, and one satellite from three stations for the rest) for each of the five examples of
 159 explosive eruptions in Japan. (bottom) High-pass filtered STEC changes over 45 minutes periods
 160 (from 15 minutes before eruption to 30 minutes after eruption) for the five explosive eruptions studied
 161 here. Small disturbances can be seen 10-15 minutes after the eruptions.

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Following Heki (2006), we try to adjust a simple function

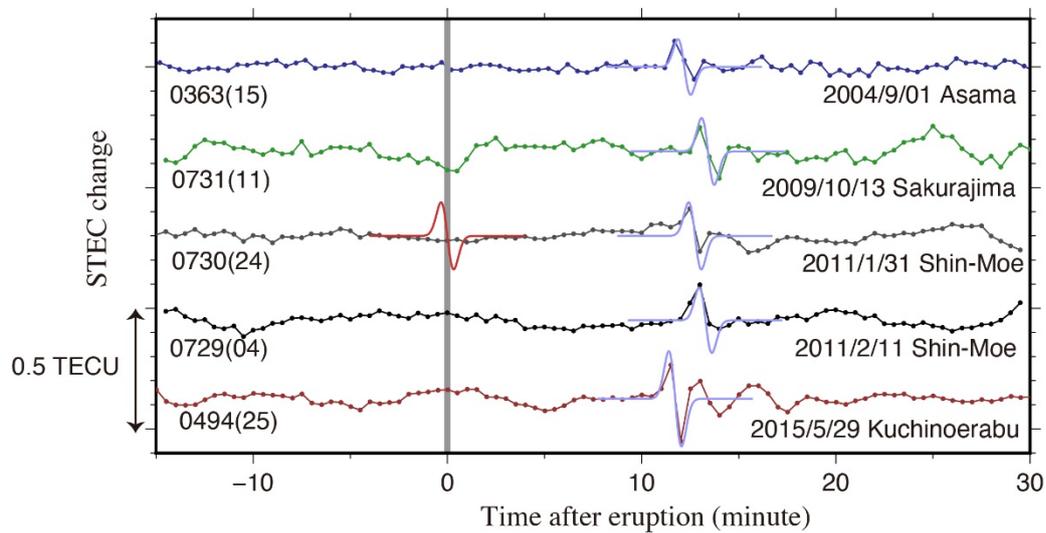
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$$f(t) = -at \exp\left(\frac{-t^2}{2\sigma^2}\right), \quad (1)$$

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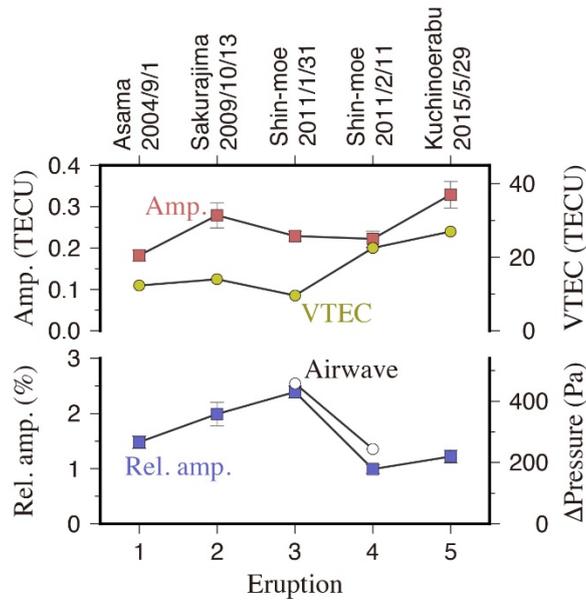
made of a set of positive and negative pulses (a curve in red in the middle of Figure 4), to the disturbances observed by five eruptions in Figure 3. This function has a maximum and minimum at $t=-\sigma$ and $t=\sigma$, respectively. The two parameters a and σ , representing the amplitude and period, respectively, were tuned to minimize the root-mean-squares (rms) of differences between the synthesized and the observed disturbances. The values $\sigma=19.5$ resulted in good fits to the majority of time series, which corresponds to 78 seconds (1.3 minutes) as the approximate period (i.e. $4 \times \sigma$) of the disturbance. From the adjusted values of a , we obtained the peak-to-peak amplitude, i.e. $f(-\sigma)-f(\sigma)$, as summarized in Figure 5. Here we do not discuss the propagation velocity of the ionospheric disturbances because they are known to propagate in the acoustic wave velocity (~ 0.8 km/s) in the ionospheric F region (Heki, 2006). The same calculation has been done for all the three examples from each eruption and the average peak-to-peak amplitudes are compared in Figure 5.



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Figure 4. Fit of the model function (shown as a red curve in the middle) to one of the STEC curves in Figure 3 for each eruption. The period of the TEC fluctuation (Heki, 2006) is fixed to 1.3 minutes, and adjusted the amplitudes and time lags to minimize the difference between the model function and the observed TEC.

Volcanic eruptions excite acoustic waves in neutral atmosphere layer and ionospheric electrons move together with such neutral atmospheric molecules. Naturally, the strength of the TEC disturbances is largely influenced by the electron density in the F region of the ionosphere. In order to study objective index for the explosion intensity, it will be reasonable to normalize the amplitudes of STEC changes with background electron densities in the F region. Here we used background VTEC to normalize such amplitudes and express the TEC amplitudes relative to them (blue squares in Figure 5). VTEC values at the time and location of eruptions are obtained from Global Ionospheric Maps (GIM; Mannucci et al. 1998).



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Figure 5. Comparison of the TECs (shown in Figure 3) in absolute (red) and relative (i.e. ration to the background VTEC) (blue) amplitudes. The yellow circles show VTEC values at the time and place of the eruptions calculated using global ionospheric models. For the two eruptions of the Shin-Moe volcano, we compare amplitudes in atmospheric pressure changes detected by the same sensor ~2.6 km from the volcano caused by airwaves by the explosions.

205 5. Discussion

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207 Intensities of the explosive volcanic eruptions can be studied by measuring amplitudes of
208 airwaves excited by the explosions as atmospheric pressure changes. However, different
209 distance of the ground sensors from the volcanoes and different topographic conditions make
210 it difficult to compare such intensities for different volcanoes. In Figure 5, we compare
211 atmospheric pressure changes by the airwaves for the the January 31 and February 11
212 explosions of the Shin-Moe volcano detected using the same sensor at the YNN station (JMA,
213 2011).

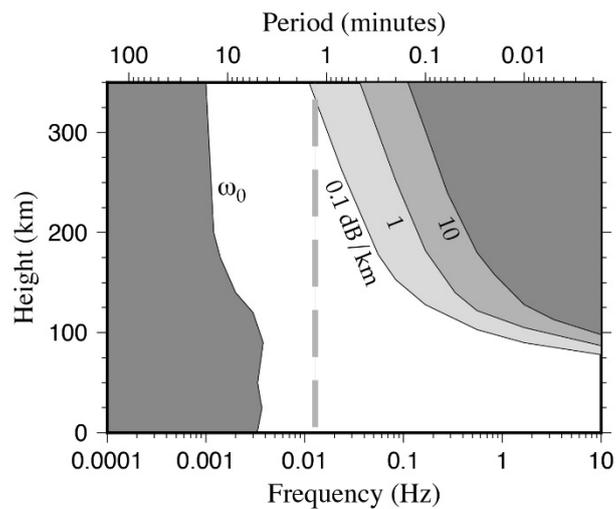
214 As seen in Figure 5, these two eruptions show similar amplitudes of STEC changes.
215 However, the background VTEC at the time of the February eruption was more than twice
216 as strong as those in the January eruption. Hence, relative amplitude of the January eruption
217 becomes twice as large as the February eruption. This is in agreement with the difference of
218 the pressure changes for these two eruptions (458.5 Pa for the January 31 eruption, and 244.3
219 Pa for the February 11 eruption).

220 This suggests the validity of using the relative amplitudes of the ionospheric STEC
221 changes as the new index to describe the intensities of volcanic explosions for different
222 volcanoes. Its benefit is that we do not rely on the deployment of infrasound sensors, i.e. we
223 can use this index whenever enough number of continuous GNSS stations are available on
224 the southern/northern side of the volcano in northern/southern hemisphere. Its drawback is
225 that this index can be used only for strong volcanic explosions occurring when number of
226 ionospheric electrons are sufficient (e.g. during daytimes). In fact, there were two explosions

227 of the Shin-Moe volcano (Feb.1 20:25 UT, and Feb. 13 20:07 UT) with stronger airwaves
228 than the February 11 02:36 UT eruption. However, we cannot find ionospheric disturbances
229 for these explosions because of small background VTEC early in the morning.

230 At last, we discuss the origin of the period of the observed TEC variations. As seen in
231 Figure 3, TEC changes by the five different volcanic explosions have similar periods of ~1.3
232 minutes. Such a uniformity suggests its origin in the atmospheric structure rather than
233 characteristics of the volcanic eruptions. In Figure 6, we compare this period with the diagram
234 of atmospheric attenuation of acoustic waves with various periods at different altitudes
235 (Blanc, 1983).

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239 **Figure 6.** Frequency (~12.8 mHz) and period (~1.3 minutes) of TEC oscillations by explosive
240 volcanic eruptions (vertical dashed line) drawn over the figure by Blanc (1983) showing the
241 attenuation of airwaves in the Earth's atmosphere. The frequency 12.8 mHz corresponds to the higher
242 end of the atmospheric bandpass filter.

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244 Figure 6 shows that 1.3 minutes corresponds to the shortest period of the airwaves that can
245 reach ionosphere without large attenuations, i.e. 12.8 mHz corresponds to the high end
246 frequency of the atmospheric band-pass filter. Infrasound records observed at ground sensors
247 associated with explosive volcanic eruptions have stronger powers in periods much shorter
248 than 1.3 minutes (e.g. Matoza et al., 2019). However, only those with periods 1.3 – 4.0
249 minutes can reach the ionospheric F region. Because the original spectrum had larger powers
250 for higher frequencies, we would have detected the 12.8 mHz component as the TEC changes
251 at the F region altitude (~300 km).

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253 Acknowledgements

254 We thank two reviewers for constructive reviews. This research was supported by the World Class
255 Professor 2019 program by the Indonesian government.

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257 References

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