

Three-dimensional distribution of ionospheric anomalies prior to three large earthquakes in Chile

Liming He^{1,2} & Kosuke Heki²

¹Institute for Geo-informatics & Digital Mine Research, College of Resources and Civil Engineering, Northeastern University, Shenyang, China

²Department of Earth and Planetary Sciences, Hokkaido University, Sapporo, Japan

Abstract

Using regional Global Positioning System (GPS) networks, we studied three-dimensional spatial structure of ionospheric total electron content (TEC) anomalies preceding three recent large earthquakes in Chile, South America, i.e. the 2010 Maule (M_w 8.8), the 2014 Iquique (M_w 8.2), and the 2015 Illapel (M_w 8.3) earthquakes. Both positive and negative TEC anomalies, with areal extent dependent on the earthquake magnitudes, appeared simultaneously 20-40 minutes before the earthquakes. For the two mid-latitude earthquakes (2010 Maule and 2015 Illapel), positive anomalies occurred to the north of the epicenters at altitudes 150-250 km. The negative anomalies occurred further to the north at higher altitudes 200-500 km. This lets the epicenter, the positive and negative anomalies align parallel with the local geomagnetic field, which is a typical structure of ionospheric anomalies occurring in response to positive surface electric charges.

1. Introduction

Ionospheric Total Electron Contents (TEC), derived by comparing phases of two L band microwave signals from Global Positioning System (GPS) satellites, allow us to study earthquakes from unique points of view. Coseismic ionospheric disturbances sometimes provide information on rupture processes of earthquakes [e.g. *Heki et al.*, 2008]. Their amplitudes depend on moment magnitudes (M_w) of earthquakes [*Cahyadi and Heki*, 2015], and ionospheric monitoring may contribute to early warning of tsunami arrivals [e.g. *Astafyeva et al.*, 2013].

Apart from these coseismic disturbances, *Heki* [2011] found ionospheric electron enhancements starting ~40 minutes before the 2011 Tohoku-oki earthquake (M_w 9.0), Japan, using a dense array of continuous GPS stations. Through debates between critical papers [*Kamogawa and Kakinami*, 2013; *Utada and Shimizu*, 2014; *Masci et al.*, 2015] and replies to them [*Heki and Enomoto*, 2013; 2014; 2015], *Heki and Enomoto* [2015] showed that the enhancements preceded eight past earthquakes with M_w 8.2 or more. They found that the enhancements started about 20/40 minutes prior to M_w 8/9 earthquakes, and that the changes in vertical TEC (VTEC) rates depend on M_w as well as background absolute VTEC. Although similar changes often occur due to geomagnetic activities, *Heki and Enomoto* [2015] demonstrated that they are not frequent enough to

47 account for the observed preseismic anomalies. The reader is referred to the introduction
48 of *Heki and Enomoto* [2015] for the history of the arguments.

49 These past papers mainly focused on the reality of enhancements and their significance
50 among space-weather origin disturbances, and little discussed spatial structures and areal
51 extents of the anomalies. Here, we study three-dimensional (3D) structures of preseismic
52 ionospheric anomalies, in order to shed light on the underlying physical processes. We
53 focus on three recent large interplate earthquakes in Chile, South America, i.e. the 2010
54 Maule (M_w 8.8), the 2014 Iquique (M_w 8.2), and the 2015 Illapel (M_w 8.3) earthquakes.
55 Large number of continuous GPS stations distributed over South America make this
56 region ideal for such studies.

57

58 **2. GPS Data and TEC analysis procedures**

59 **2.1 The three Chilean earthquakes**

60 In this study, we analyzed the behaviors of ionospheric TEC before and after the three
61 large earthquakes in Chile using regional GPS data (Figure 1). The 2010 February 27
62 Maule earthquake ruptured the boundary between the Nazca and the South America
63 Plates known as the Constitución-Concepción seismic gap in Central Chile [*Madariaga*
64 *et al.*, 2010]. The 2014 April 1 Iquique earthquake ruptured the same plate boundary
65 around the Peru-Chile border [*Ruiz et al.*, 2014]. The 2015 September 16 Illapel
66 earthquake occurred in a segment ~500 km north of the Maule earthquake [*Ye et al.*,
67 2015]. Geographic latitudes of their epicenters are 36.1S, 19.6S, and 31.6S, respectively,
68 and their geomagnetic latitudes are lower by ~10 degrees. The Maule earthquake
69 occurred past midnight (03:34:14) in local time (LT) and the other two occurred in the
70 evening, i.e. at 20:46:47 LT (2014 Iquique) and at 19:54:33 LT (2015 Illapel).

71

72 **2.2 GPS-TEC data processing**

73 We first obtained slant TEC (STEC), number of electrons integrated along the travel
74 path of microwave signals, and removed the phase ambiguities by letting the TEC time
75 series derived by carrier phases align with those by pseudo-ranges [*Calais and Minster*,
76 1995; *Mannucci et al.*, 1998]. The STEC data still include satellite and receiver inter-
77 frequency biases (IFBs). We obtained the satellite IFBs from the header information of
78 global ionospheric maps available from University of Berne, Switzerland [*Schaer et al.*,
79 1998]. We inferred the receiver IFBs by minimizing the scatter of nighttime VTEC at
80 individual stations [*Rideout and Coster*, 2006].

81 After removing IFB, we calculated the absolute VTEC by multiplying the STEC with
82 the cosine of the incident angle of the line-of-sight with a thin shell at 300 km above the
83 ground (In drawing SIPs in Figures 2 and 3, we assumed different heights). We use
84 VTEC throughout this study because they are free from apparent variations due to
85 changing satellite elevations. Geomagnetic activities were low when the 2010 Maule and
86 the 2014 Iquique earthquakes occurred, but was moderately high during the 2015 Illapel
87 earthquake (Figure S1).

88 We followed *Heki and Enomoto* [2015] to identify bends (breaks) in VTEC before
89 earthquakes using the Akaike's information criterion (AIC) (Figure S2). Figure 2 shows
90 VTEC time series observed using different pairs of GPS satellites and receivers over 2-3
91 hour intervals including the three Chilean earthquakes. We modeled the VTEC curves
92 with the polynomials of time with degree 3-5, excluding the intervals from the onsets of

93 the anomalies detected using AIC to 20 minutes after the earthquakes. We define positive
94 and negative departures from the models as anomalous increases (enhancements) and
95 decreases of VTEC, respectively.

96 Some stations show clear coseismic disturbances, e.g. CRZL-PRN20, AREQ-PRN01
97 and TILC-PRN14, in Figure 2a, b and c, respectively. Such disturbances clearly appear
98 where SIPs and stations are on the northern side of the epicenter, owing to interaction
99 with geomagnetic fields [Rolland *et al.*, 2013]. Coseismic electron depletions are clear
100 above the regions of large vertical coseismic crustal movements [Shinagawa *et al.*, 2013],
101 e.g. RCSD-PRN23, LYAR-PRN23 and CMPN-PRN24, in Figure 2a, b and c,
102 respectively, but are smaller outside such regions.

103 Preseismic TEC enhancements emerge ~40 minutes before the 2010 Maule earthquake
104 (Figure 2a-2, upper half). We found that TEC starts to decrease (Figure 2a-2, lower half)
105 simultaneously at stations farther to the north of the epicenter. The 2014 Iquique (Figure
106 2b-2) and 2015 Illapel (Figure 2c-2) earthquakes also showed similar sets of VTEC
107 enhancements and decreases starting ~20 minutes before earthquakes.
108

109 **3. Preseismic TEC anomalies of the three Chilean earthquakes**

110 Figure 3 shows map distributions of the TEC anomalies at three time epochs, i.e. 30
111 minutes, 15 minutes, and immediately (30 seconds) before the 2015 Illapel earthquake.
112 The dots represent the SIP positions calculated assuming the ionospheric heights of ~170
113 km and ~420 km, for positive and negative TEC anomalies, respectively. SIP coordinates
114 depend on the assumed height of an ionospheric anomaly, and multiple SIPs obtained
115 with different satellite-station pairs are expected to converge when the assumed anomaly
116 height is correct (Figure S3). In order to constrain the altitudes of the observed positive
117 and negative anomalies, we tuned their altitudes so that they minimize the angular
118 standard deviations of the SIPs of positive and negative groups, respectively (Figure S4).

119 No anomalies exist 30 minutes before the earthquake (Figure 3a, d). Both positive and
120 negative VTEC anomalies have already emerged to the north of the epicenter ~15
121 minutes before the earthquake (Figure 3b, e). They become the largest immediately
122 before the earthquake (Figure 3c, f). Such positive and negative TEC anomalies also
123 preceded the 2010 Maule (Figure S5) and the 2014 Iquique (Figure S6) earthquakes. For
124 the 2010 and 2015 events, the positive anomalies are located just to the north of the
125 epicenters, and the negative anomalies appeared far to the north over a larger area. We
126 find fewer anomalies to the south of the epicenter. For the 2014 event, the positive
127 anomalies emerged just above the epicenter, and the negative anomalies appeared both on
128 the northern and southern sides.

129 Figure 4 compares the map views of preseismic positive VTEC anomalies immediately
130 before the three Chilean earthquakes. We tuned the ionospheric heights to minimize the
131 scatters of the positive anomalies (Figure S4). The dimensions of areas of positive VTEC
132 anomalies depend on M_w (hence on fault size); the anomaly of the 2010 M_w 8.8 Maule
133 earthquake are larger than the other two (M_w 8.2 and 8.3) earthquakes. The background
134 VTEC, on the other hand, does not seem to influence the dimension; the background
135 VTEC of the 2014 Iquique earthquake was >10 times as large as that of the 2010 Maule
136 earthquake (see Fig. 1 of Heki and Enomoto [2015]).
137

138 **4. Discussions**

139 **4.1 Onset time and the VTEC rate changes**

140 Before discussing 3D spatial distribution of the ionospheric anomalies, we briefly
141 discuss other aspects of the anomalies. Using multiple pairs of satellite and stations, we
142 found that the onset time of the anomaly was ~40, ~23, and ~22 minutes before the 2010,
143 2014, and 2015 earthquakes (Figure S2). This is consistent with other earthquakes (Fig.
144 5a in *Heki and Enomoto* [2015]).

145 *Heki and Enomoto* [2015] have already used the 2010 Maule and 2014 Iquique
146 earthquakes to derive the empirical relationship (equation 5) between the background
147 VTEC, VTEC rate changes, and M_w . This equation assumes that magnitudes of very
148 large earthquakes are already determined in the nucleation stage. The observations
149 suggest that cascading-up would not much exceed the difference between the predicted
150 and real magnitudes (0.28 in *Heki and Enomoto* [2015]). For the 2015 Illapel event, a
151 new earthquake, the background VTEC of 22 TECU and the observed break of 4.3
152 TECU/h (CMPN- PRN24) predicts M_w of 8.7. This is 0.4 larger than the actual M_w .
153 Inclusion of this new event, and future large earthquakes, would further refine the
154 coefficients of the equation, and improve the accuracy of the expected M_w .

155 **4.2 Spatial structures of preseismic ionospheric anomalies**

156 For studying 3D spatial distribution of the TEC anomalies, we use the data of the 2015
157 Illapel earthquake, for which the station distribution is the most suitable (Figure 1). The
158 conventional way is to map horizontal distribution of TEC anomalies, like Figure 3, as if
159 they occurred on a horizontal plane at a certain height. In Figure 5a, we drew a
160 “longitudinal profile”, where we plot the calculated heights and latitudes of the
161 intersections of the line-of-sight vectors with the 70W meridional plane. Because the
162 line-of-sights need to intersect with the plane at high angles, we used only PRN14 and 25.
163 The profile shows that the positive and negative anomalies line up along the magnetic
164 field (inclination -32° [*Thebault et al.*, 2015]) from the bottom of ionosphere (~85 km
165 high) above the epicenter.
166

167 Figures 3 and 5a suggest the 3D spatial structure of the preseismic ionospheric
168 anomalies as illustrated in Figure 5b. This resembles to the numerical calculation results
169 of the ionospheric response to the positive surface electric charges (see Figure 12e of *Kuo*
170 *et al.* [2014]). For the 2010 Maule earthquake, the distribution of GPS stations (Figure 1a)
171 was not good enough to plot the longitudinal profiles, but the SIP maps (Figure S5) do
172 suggest a similar 3D structure. In the 2014 Iquique earthquake, the negative anomalies
173 exist on both the north and south of the positive anomalies (Figure S6). This difference
174 might reflect the lower geomagnetic latitude (~10S) of the 2014 earthquake epicenter.
175 We are also preparing for a systematic search for the mirror image anomalies that are
176 expected to emerge at geomagnetic conjugate points.
177

178 **4.3 Physical process responsible for preseismic TEC anomalies**

179 *Kuo et al.* [2014] demonstrated, with numerical simulation, that the westward electric
180 field in the ionosphere originated from the upward atmospheric electric current causes
181 obliquely downward $\mathbf{E} \times \mathbf{B}$ drift of electrons. This drift causes the increase and decrease of
182 electron density at altitudes of ~200 and ~400 km, respectively, and results in the positive
183 and negative anomalies lying along the geomagnetic field. The observed anomalies
184 before the 2015 Illapel earthquake (Figure 5) are consistent with this picture. The

185 nighttime ionosphere before this earthquake changed by ~10% of the background VTEC.
186 This requires the maximum density of ~10 nA m⁻² of upward atmospheric electric current
187 at the geomagnetic latitude (21.7S) of the 2015 Illapel earthquake [Kuo *et al.*, 2014].

188 A candidate mechanism to explain such currents is the outflow of positive holes from
189 fast stressed rock to unstressed rock observed in laboratory experiments [e.g. Freund *et*
190 *al.*, 2009]. Such currents sharply increase immediately before the failure of rock samples,
191 and then decrease exponentially over a short time after the failure. This resembles the
192 observed VTEC anomaly behaviors. Although there are no decisive evidences, the
193 present observations support the scenario that positive charges from rocks under near-
194 failure stress, possibly in the earthquake nucleation stage, cause ionospheric anomalies
195 immediately before large earthquakes.

196

197 **Acknowledgments**

198 We thank two reviewers for constructive comments. LH was supported by the China Scholarship
199 Council (CSC) and by the National Natural Science Foundation of China (grant no. 41104104).
200 The Chilean GNSS data for the 2010 Maule earthquake were provided by C. Vigny (ENS). We
201 downloaded the Argentine (RAMSAC/IGNA) and Brazilian (RBMC/IBGE) GNSS data from
202 their official webpages. We downloaded additional data from IGS (www.igs.org) and UNAVCO
203 (www.unavco.org). We thank C. L. Kuo, National Central University, Taiwan, for validating the
204 observation results using the simulation model developed in his group.

205

206 **References**

- 207 Astafyeva, E. *et al.* (2013), Parameters of seismic source as deduced from 1 Hz
208 ionospheric GPS data: Case study of the 2011 Tohoku-oki event, *J. Geophys. Res.*
209 *Space Phys.*, *118*, 5942-5950.
- 210 Calais, E., and J. B. Minster (1995), GPS detection of ionospheric perturbations
211 following the January 17, 1994, Northridge earthquake, *Geophys. Res. Lett.*, *22*, 1045-
212 1048.
- 213 Cahyadi, M.N. and K. Heki (2015), Coseismic ionospheric disturbance of the large
214 strike-slip earthquakes in North Sumatra in 2012: M_w dependence of the disturbance
215 amplitudes, *Geophys. J. Int.*, *200*, 116-129.
- 216 Freund, F.T., I.G. Kulahci, G. Cyr, J. Ling, M. Winnick, J. Tregloan-Reed, and M. M.
217 Freund (2009), Air ionization at rock surfaces and pre-earthquake signals, *J. Atmos. Sol.*
218 *Terr. Phys.*, *71*, 1824-1834, doi:10.1016/j.jastp.2009.07.013.
- 219 Heki, K. (2011), Ionospheric electron enhancement preceding the 2011 Tohoku-Oki
220 earthquake, *Geophys. Res. Lett.*, *38*, L17312.
- 221 Heki, K., and Y. Enomoto (2013), Preseismic ionospheric electron enhancements
222 revisited, *J. Geophys. Res. Space Phys.*, *118*, 6618-6626.
- 223 Heki, K. and Y. Enomoto (2014), Reply to comment by K. Heki and Y. Enomoto on
224 "Preseismic ionospheric electron enhancements revisited", *J. Geophys. Res. Space Phys.*,
225 *119*, doi:10.1002/2014JA020223.
- 226 Heki, K., and Y. Enomoto (2015), M_w dependence of the preseismic ionospheric electron
227 enhancements, *J. Geophys. Res. Space Phys.*, *120*, 7006-7020.
- 228 Heki, K., Y. Otsuka, N. Choosakul, N. Hemmakorn, T. Komolmis, and T. Maruyama
229 (2006), Detection of ruptures of Andaman fault segments in the 2004 Great Sumatra

230 Earthquake with coseismic ionospheric disturbances, *J. Geophys. Res. Solid Earth*, *111*,
231 B09313, doi:10.1029/2005JB004202.

232 Kamogawa, M. and Y. Kakinami (2013), Is an ionospheric electron enhancement
233 preceding the 2011 Tohoku-oki earthquake a precursor?, *J. Geophys. Res. Space*
234 *Phys.*, *118*, 1-4, doi:10.1002/jgra.50118.

235 Kuo, C. L., L. C. Lee, and J. D. Huba (2014), An improved coupling model for the
236 lithosphere-atmosphere-ionosphere system, *J. Geophys. Res. Space Phys.*, *119*, 3189-
237 3205, doi:10.1002/2013JA019392.

238 Madariaga, R., M. Métois, C. Vigny, and J. Campos (2010), Central Chile finally breaks,
239 *Science*, *328*, 181-182.

240 Mannucci, A., B. Wilson, D. Yuan, C. Ho, U. Lindqwister, and T. Runge (1998), A
241 global mapping technique for GPS-derived ionospheric total electron content
242 measurements, *Radio Sci.*, *33*, 565-582.

243 Masci, F., J. N. Thomas, F. Villani, J. A. Secan, and N. Rivera (2015), On the onset of
244 ionospheric precursors 40 min before strong earthquakes, *J. Geophys. Res. Space*
245 *Phys.* *120*, 1383-1393, doi:10.1002/2014JA020822.

246 Rideout, W., and A. Coster (2006), Automated GPS processing for global total electron
247 content data, *GPS Solut.*, *10*, 219-228.

248 Rolland, L. M., M. Vergnolle, J. M. Nocquet, A. Sladen, J. X. Dessa, F. Tavakoli, H. R.
249 Nankali, and F. Cappa (2013), Discriminating the tectonic and non-tectonic
250 contributions in the ionospheric signature of the 2011, Mw7. 1, dip-slip Van
251 earthquake, Eastern Turkey, *Geophys. Res. Lett.*, *40*, 2518-2522.

252 Ruiz, S., M. Metois, A. Fuenzalida, J. Ruiz, F. Leyton, R. Grandin, C. Vigny, R.
253 Madariaga, and J. Campos (2014), Intense foreshocks and a slow slip event preceded
254 the 2014 Iquique Mw 8.1 earthquake, *Science*, *345*, 1165-1169.

255 Schaer, S., W. Gurtner, and J. Feltens (1998), IONEX: The ionosphere map exchange
256 format version 1, in *Proceedings of the IGS AC workshop*, Darmstadt, Germany.

257 Shinagawa, H., T. Tsugawa, M. Matsumura, T. Iyemori, A. Saito, T. Maruyama, H. Jin,
258 M. Nishioka, and Y. Otsuka (2013), Two-dimensional simulation of ionospheric
259 variations in the vicinity of the epicenter of the Tohoku-Oki earthquake on 11 March
260 2011, *Geophys. Res. Lett.*, *40*, 5009-5013.

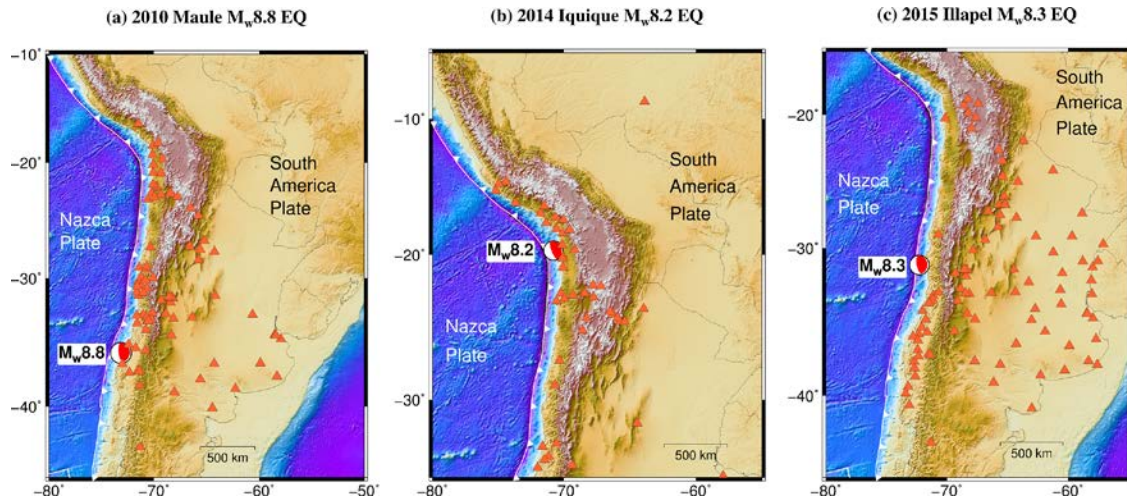
261 Thebault, E. et al. (2015), International Geomagnetic Reference Field: the 12th generation,
262 *Earth Planets Space*, *67*, 1-19.

263 Utada, H., and H. Shimizu (2014), Comment on “Preseismic ionospheric electron
264 enhancements revisited” by K. Heki and Y. Enomoto, *J. Geophys. Res. Space Phys.*,
265 *119*, doi:10.1002/2014JA020044.

266 Ye, L., T. Lay, H. Kanamori, and K. D. Koper (2016), Rapidly estimated seismic source
267 parameters for the 16 September 2015 Illapel, Chile Mw 8.3 earthquake, *Pure Appl.*
268 *Geophys.*, *173*, 321-332.

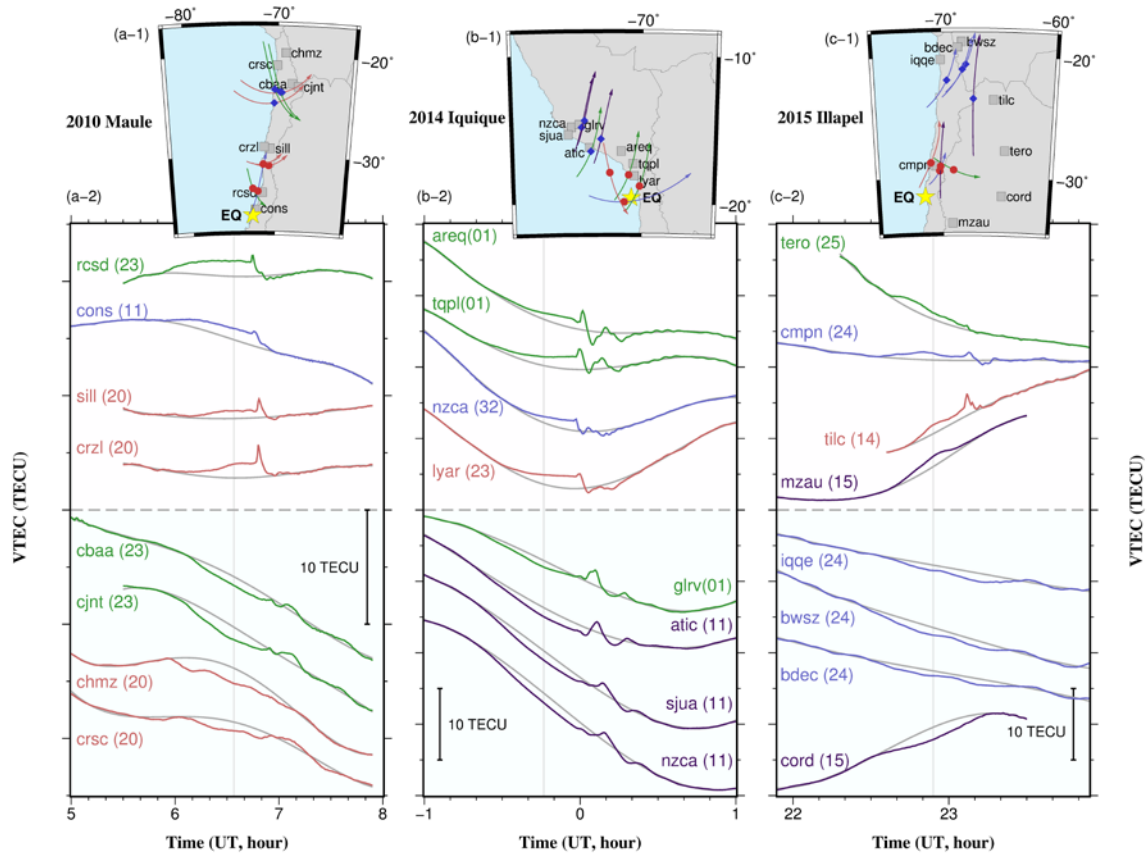
269

270 **Figures**
271



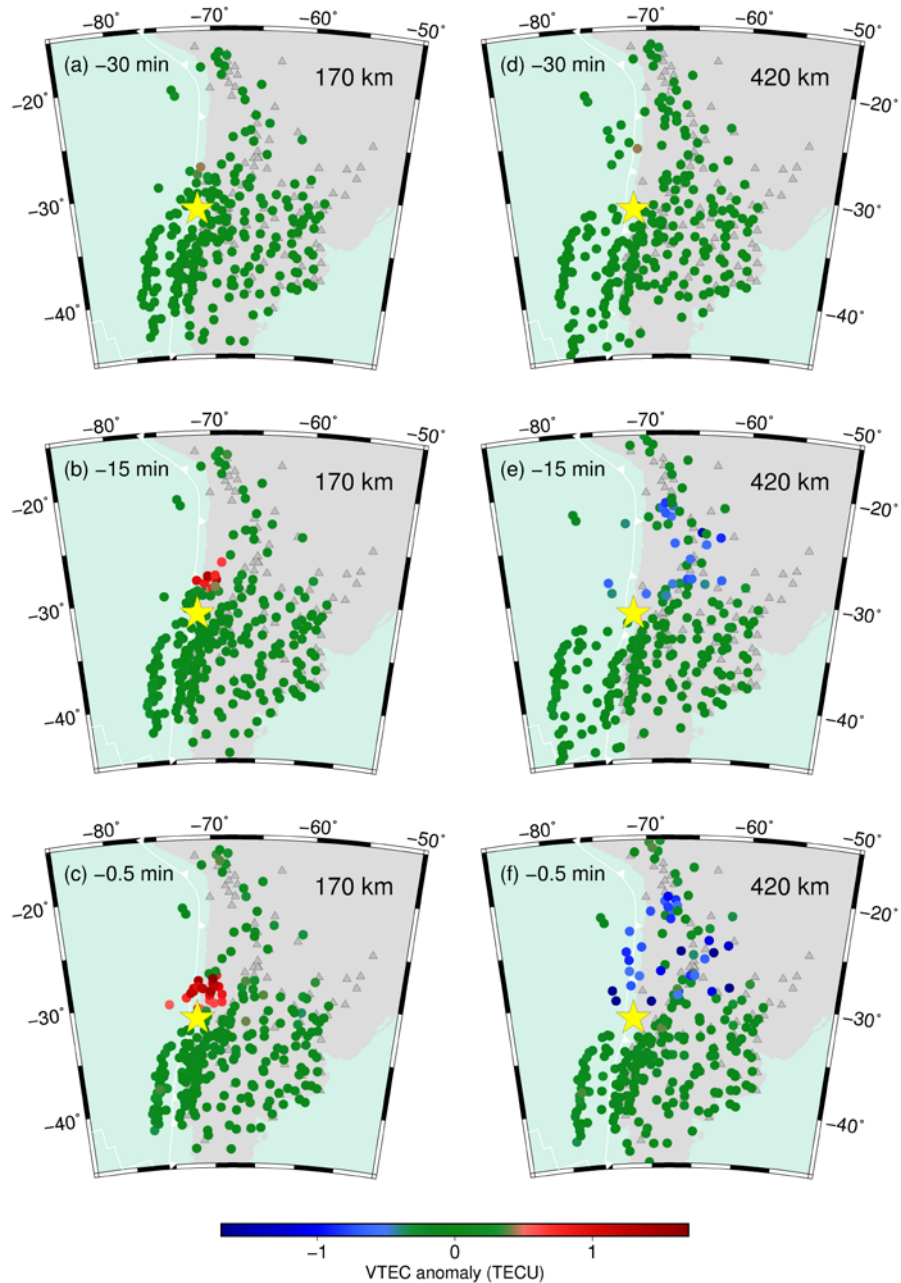
272
273
274
275
276
277
278

Figure 1. GPS stations (red triangles) used to study the three Chilean earthquakes, (a) 83 stations for the 2010 Maule, (b) 47 stations for the 2014 Iquique, and (c) 91 stations for the 2015 Illapel earthquakes. The beach balls show the epicenters and focal mechanisms.



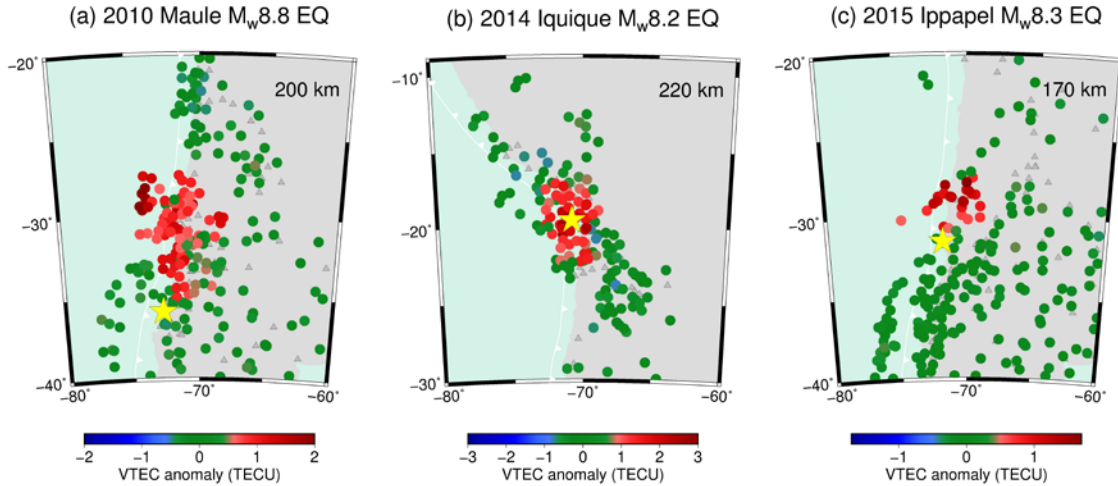
280
 281
 282
 283
 284
 285
 286
 287
 288
 289
 290
 291

Figure 2. The VTEC time series observed with eight pairs of station-satellite (same colors for same satellites) showing preseismic enhancements and decreases shown in the upper and lower halves, respectively. The grey curves are the reference models, from which we define VTEC anomalies shown in Figure 3. The vertical grey lines indicate earthquake occurrence times. The maps at the top show the positions of GPS stations (gray squares) and the SIP trajectories (red dots and blue diamonds indicate the earthquake times for stations showing enhancements and decreases, respectively) over the studied intervals. We assumed 200 and 400 km for ionospheric heights in drawing SIP tracks for enhancements and decreases, respectively. The yellow stars show the epicenters.



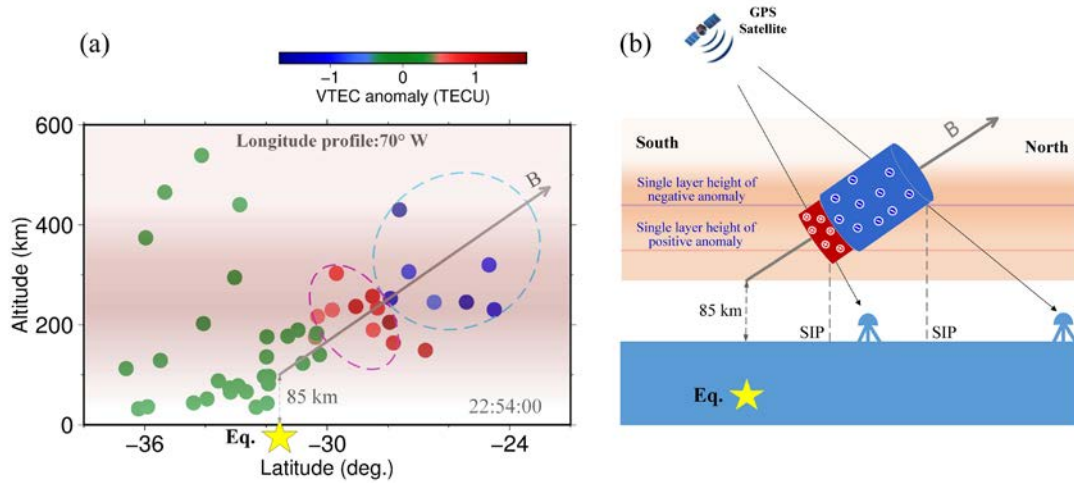
292
 293
 294
 295
 296
 297
 298
 299
 300

Figure 3. Distribution of SIPs showing preseismic positive/negative VTEC anomalies at (a/d) 30 minutes, (b/e) 15 minutes, and (c/f) 0.5 minutes before the 2015 Illapel earthquake. We used five GPS satellites (PRN12, 14, 15, 24 and 25). We derived the SIP positions in (a-c) and (d-f) assuming the ionospheric heights of 170 and 420 km, respectively. The yellow star shows the epicenter, and gray triangles indicate GPS receivers. We removed negative (< -0.5 TECU) and positive ($> +0.5$ TECU) anomalies from (a-c) and (d-f), respectively, for visual clarity.



301
 302
 303
 304
 305
 306
 307
 308

Figure 4. Preseismic VTEC enhancements 0.5 min before the 2010 Maule (a), 2014 Iquique (b), and 2015 Illapel (c) earthquakes. We assumed the ionospheric altitudes of 200, 220, and 170 km in calculating SIP positions for the three earthquakes, respectively. The yellow stars show the epicenters. The gray triangles represent the positions of GPS receivers.



309
 310
 311
 312
 313
 314
 315
 316
 317

Figure 5. (a) Longitudinal profile at 70°W of VTEC anomalies immediately before the 2015 Illapel earthquake drawn using PRN 14 and 25. We show latitudinal profile in Figure S7. (b) Schematic illustration showing the 3D distribution of positive and negative anomalies. A thick grey arrow shows the geomagnetic field. The red and the blue regions show the positive and negative electron density anomalies. The yellow stars show the epicenter.

318



319

320

Geophysical Research Letters

321

Supporting Information for

322

Three-dimensional distribution of ionospheric anomalies prior to three large earthquakes in Chile

323

324

325

Liming He^{1,2} & Kosuke Heki²

326

327

¹Institute for Geo-informatics & Digital Mine Research, College of Resources and Civil Engineering, Northeastern University, Shenyang, China

328

329

²Department of Earth and Planetary Sciences, Hokkaido University, Sapporo, Japan

330

331

Contents of this file

332

333

Figures S1 to S6

334

335

Additional Supporting Information (Files uploaded separately)

336

337

None

338

339

340

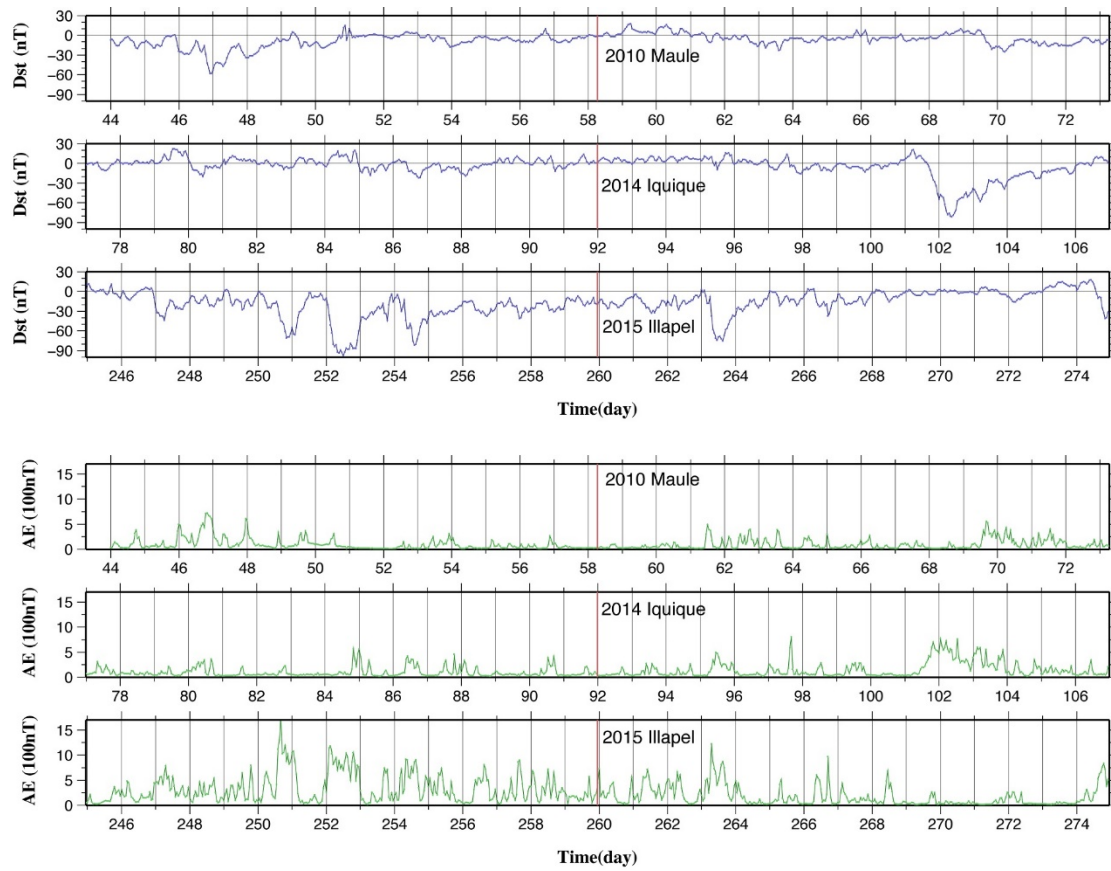
Introduction

341

This supporting information provides seven figures: time series of geomagnetic indices 15 days before and after the earthquake days (Figure S1), inference of the onset times of VTEC anomalies using AIC (Figure S2), concept of determination of ionospheric anomaly heights (Figure S3, S4), the evolution of positive and negative ionospheric anomalies before the 2010 Maule earthquake (Figure S5), and before the 2014 Iquique earthquake (Figure S6), and the latitudinal profile of VTEC anomalies (Figure S7).

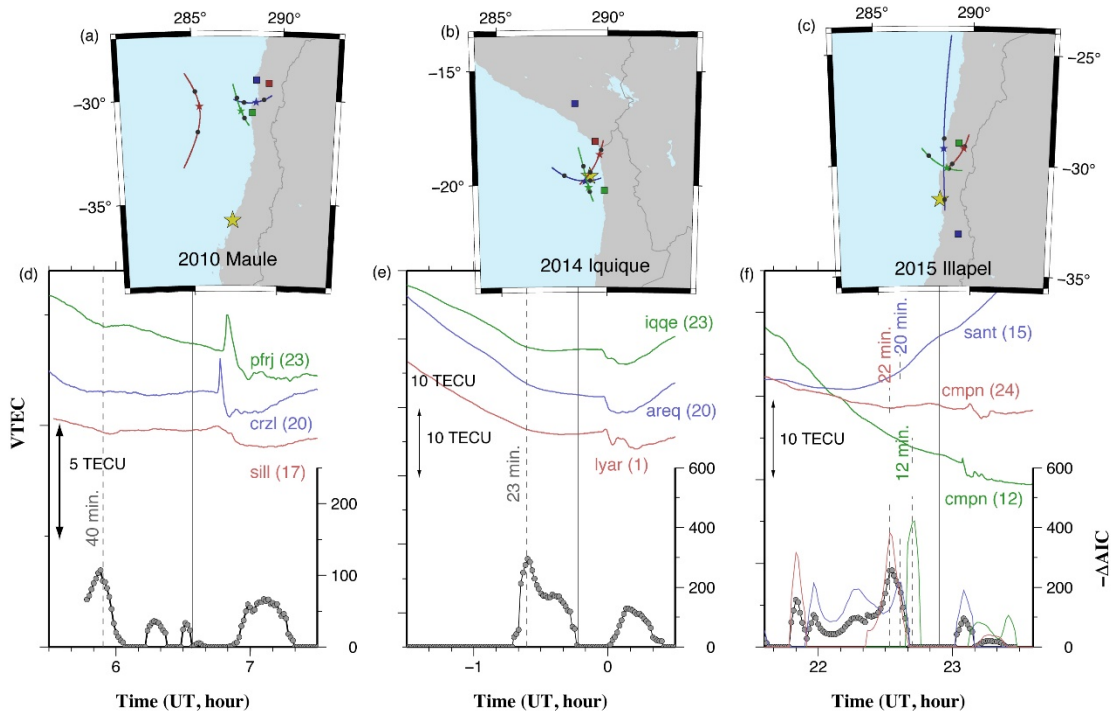
346

347



348

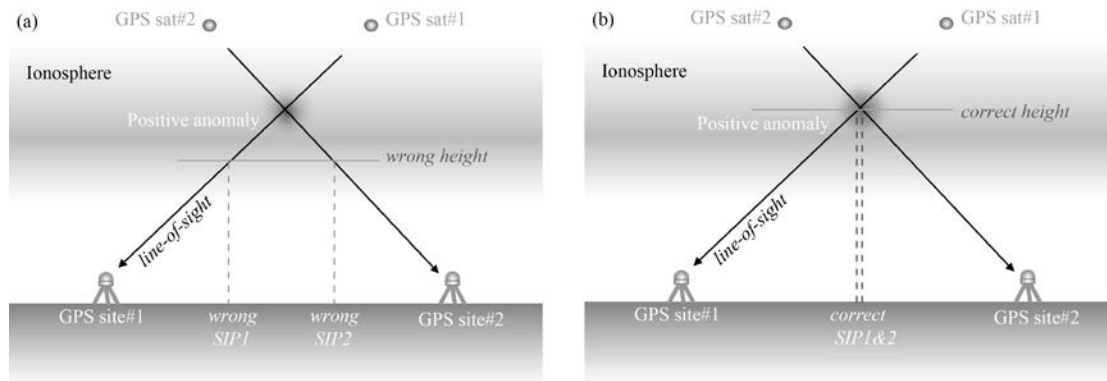
349 **Figure S1.** Time series of the *Dst* and AE indices ± 15 days before and after the occurrence days
 350 of the 2010 Maule, the 2014 Iquique, and the 2015 Illapel earthquakes. Data taken from
 351 NASA/GSFC Omniweb (<http://omniweb.gsfc.nasa.gov/form/dx1.html>).
 352



353

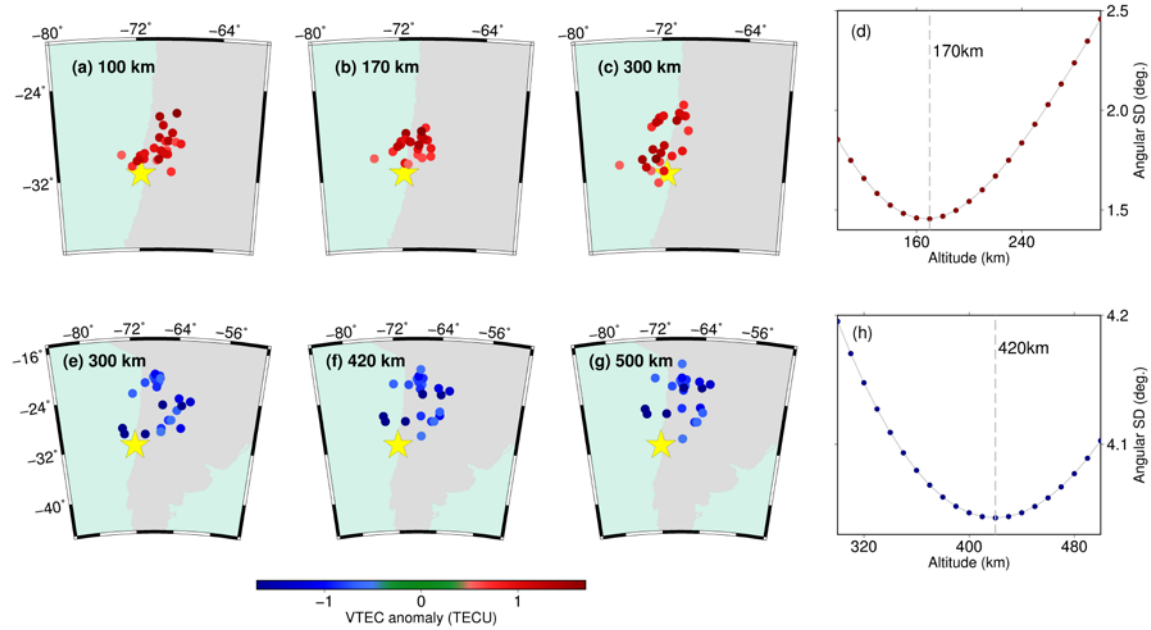
354 **Figure S2.** The onset times of preseismic VTEC anomalies inferred using Akaike's information
 355 criterion (AIC) following the method of *Heki and Enomoto* [2015]. The colored lines in (d-f)
 356 show three selected VTEC curves before and after the earthquake, and the vertical lines indicate
 357 earthquake occurrence times. In calculating $-\Delta AIC$, we used moving time windows of ± 15
 358 minutes, and the absolute and relative thresholds of 1 TECU/h and 50 %, respectively. The thick
 359 grey curves show the $-\Delta AIC$ time series. They are the averages of the $-\Delta AIC$ changes of the
 360 three time series for the 2010 (d) and 2014 (e) earthquakes. For the 2015 earthquake (f), only
 361 those of the two (CMPN-PRN24, SANT-PRN15) are averaged considering that the break in the
 362 CMPN-PRN12 curve shows the entry of the line-of-sight into the positive anomaly rather than its
 363 onset. Significant positive breaks can be seen ~ 40 min, ~ 23 min and ~ 22 min before the 2010
 364 Maule ($M_w 8.8$), the 2014 Iquique ($M_w 8.2$), and the 2015 Illapel ($M_w 8.3$) earthquakes, respectively.
 365 $-\Delta AIC$ showed slightly different peak times in the 2015 case, and those for the individual
 366 satellite-station pairs are shown with thin colored curves in (f). The upper right figure (a-c) shows
 367 the locations of GPS stations (squares), epicenters (yellow stars) and SIP traces (1 hour time
 368 marks shown with dots, and earthquake occurrence times are shown with stars). Their colors (red,
 369 blue, green) correspond to those in (d-f). We assumed ionospheric height of 200 km to obtain the
 370 SIP traces.

371



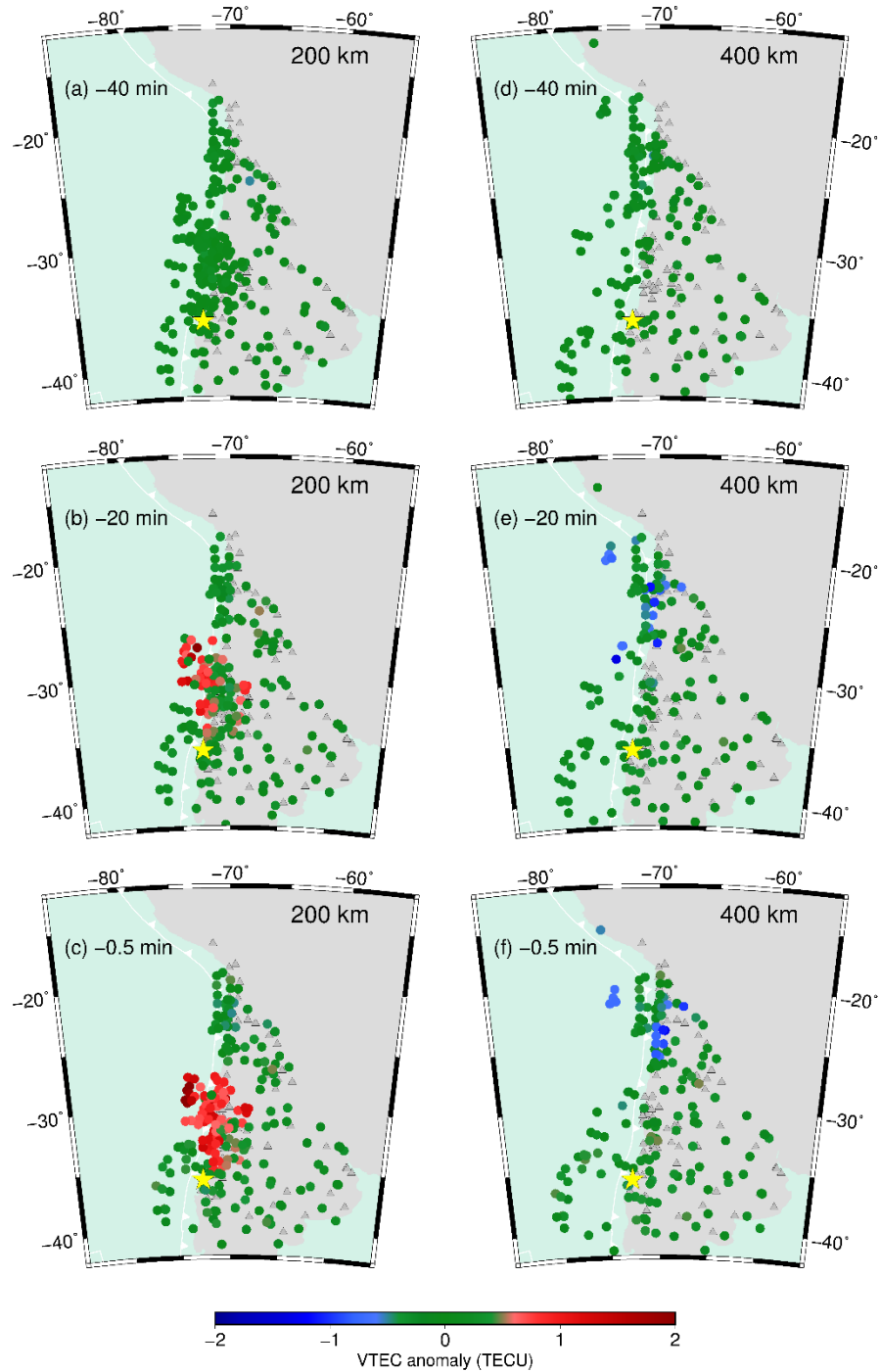
372

373 **Figure S3.** Two satellite-site pairs (sat#1-site#1 and sat#2-site#2) show positive TEC anomalies
 374 because their line-of-sights penetrate the ionospheric positive anomaly. With a wrong assumption
 375 of the anomaly height, SIP positions of the two satellite-receiver pairs would deviate from each
 376 other (a). On the other hand, the correct height would let their SIP positions converge (b). We
 377 separately determined the approximate heights of the positive and negative ionospheric anomalies,
 378 by minimizing the scatters of SIP positions of satellite-station pairs showing positive and negative
 379 TEC anomalies. These heights are used in drawing Figures 3, 4, S5 and S6.



380
 381 Figure S4. We determine the altitudes of preseismic electron density increases and decreases
 382 by minimizing the angular standard deviations (SD) of the SIPs showing positive and
 383 negative anomalies. The SIPs of positive anomalies immediately before the 2015 Illapel
 384 earthquake (22:54 UT), obtained by PRN12, 14, 15, 24 and 25, are drawn assuming the
 385 height of thin ionosphere of (a) 100 km, (b) 170 km, and (c) 300 km. In (d) we show the
 386 behavior of angular SD when we changed the height from 100 km to 300 km with the step of
 387 10 km. The angular SD showed clear minimum at 170 km. We repeated the same procedure
 388 for the negative anomalies obtained by PRN12, 15, 24 and 25, assuming the height of thin
 389 ionosphere of (e) 300 km, (f) 420 km, and (g) 500 km. The angular SD showed the minimum
 390 at 420 km (h).

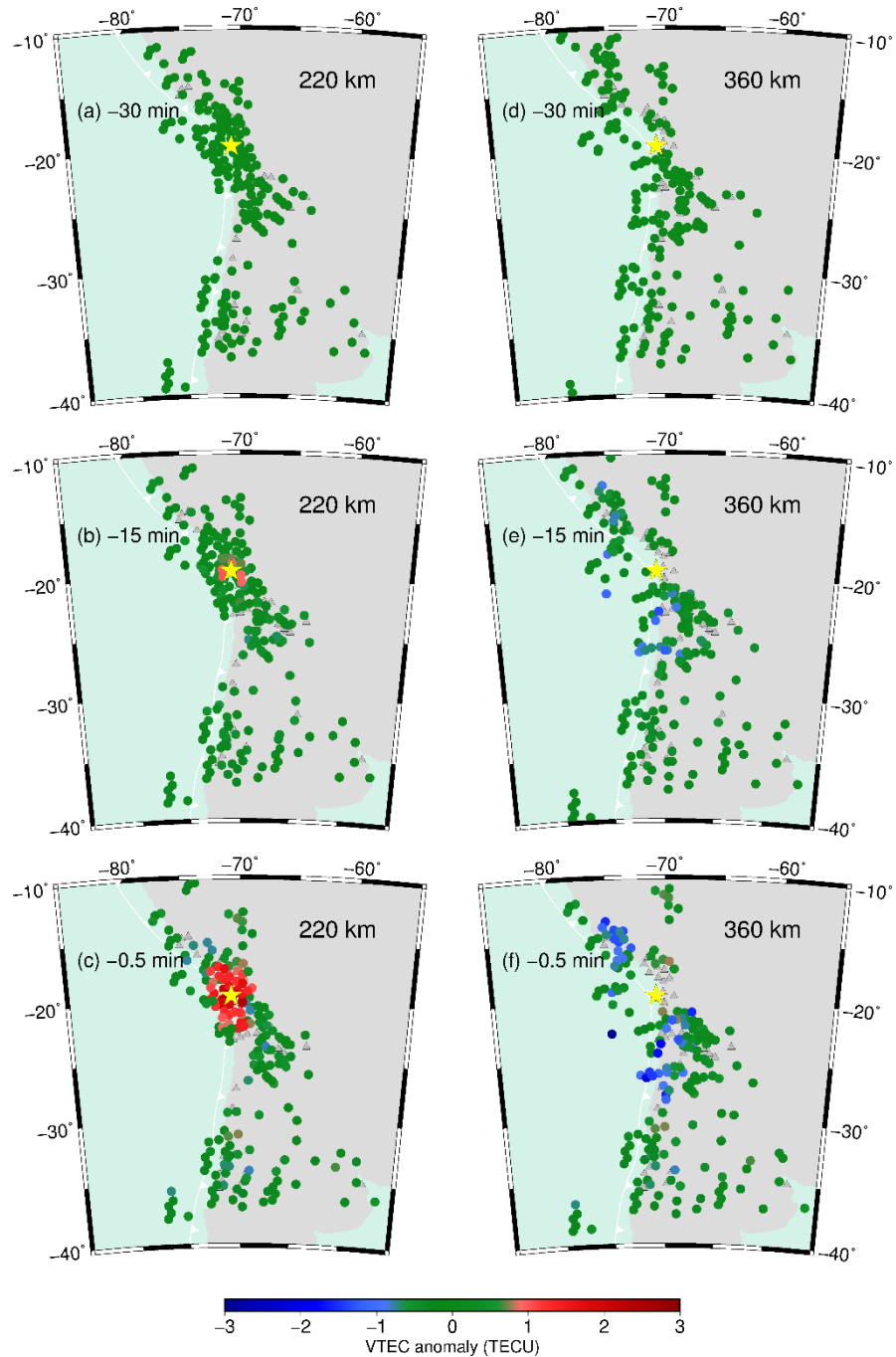
391
 392



393

394 **Figure S5.** Distribution of positive (a-c) and negative (d-f) VTEC anomalies before the 2010 M_w 8.8
 395 Maule earthquake. We used five GPS satellites (PRN11, 13, 17, 20, and 23). We assumed that the
 396 positive and negative anomalies exist at heights of 200 km and 400 km, respectively. For the detail, see
 397 the caption of Figure 3.

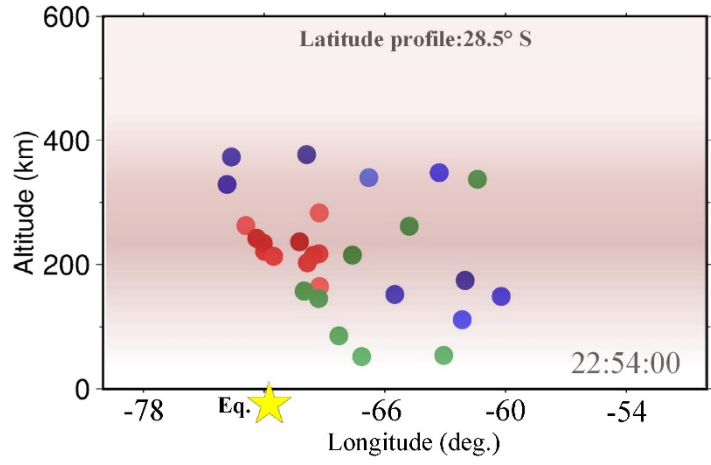
398



399

400 **Figure S6.** Distribution of positive (a-c) and negative (d-f) VTEC anomalies before the 2014 M_w 8.2
 401 Illapel earthquake. We used five GPS satellites (PRN01, 11, 13, 20, and 23). We assumed that the
 402 positive and negative anomalies exist at heights of 220 km and 360 km, respectively. For the detail, see
 403 the caption of Figure 3.

404



405

406 **Figure S7.** Latitudinal profiles of VTEC anomalies at 22:54:00, 30 seconds before the 2015 Illapel
 407 earthquake, using PRN 15, drawn assuming the latitudes of the profile at 28.5°S, ~300 km to the north of
 408 the epicenter. We can see that intersections showing positive anomalies concentrate at height of 170-250
 409 km.

410